

Regularization Parameter Estimation: Stabilization of LSQR Algorithms by Iterative Reweighting for Inversion of $3D$ Gravity Data

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Motivation: Large Scale Gravity Inversion

Parameter estimation on the projected problem

UPRE is a good estimator [RVA15]

Identifying the weight parameter in the GCV [CNO08]

Identifying the optimal Subspace

Appearance of Noise in the Subspace [HPS09]

Minimization of the GCV for the truncated SVD [CKO15]

Simulations: Two dimensional Examples

Iteratively Reweighted Regularization [LK83]

Inversion of undersampled gravity data

Conclusions

Motivation: 3D Gravity Inversion

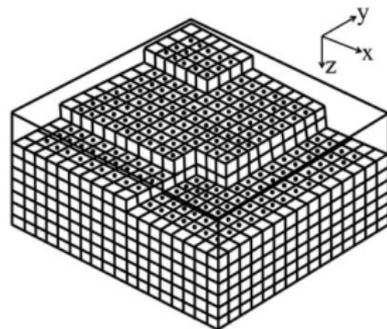
Observation point $\mathbf{r} = (x, y, z)$

Vertical gravitational attraction $g(\mathbf{r})$

$$g(\mathbf{r}) = \Gamma \int_{d\Omega} \varrho(\mathbf{r}') \frac{\mathbf{r}' - \mathbf{r}}{|\mathbf{r}' - \mathbf{r}|^3} d\Omega'$$

Density $\varrho(\mathbf{r}')$ at $\mathbf{r}' = (x', y', z')$

Newton gravitational constant: Γ



Aim: Given surface observations g_{ij} find volume density ϱ_{ijk}

Gravity Measurements g_{ij} \mathbf{b} on surface at m cells.

Density ρ_{ijk} \mathbf{x} on volume of n cells.

Projection matrix $A \in \mathbb{R}^{m \times n}$

Linear System $A\mathbf{x} \approx \mathbf{b}$:

- ▶ Severely Underdetermined: $m \ll n$
- ▶ Noise contamination $\mathbf{b} = \mathbf{b}_{\text{true}} + \boldsymbol{\eta}$
- ▶ Ill-posed: $\text{cond}(A)$ large
- ▶ Relatively Large: e.g. $m = 4588$, $n = 100936$

Tikhonov Regularization:

$$\mathbf{x}(\lambda) = \underset{\mathbf{x} \in \mathbb{R}^n}{\text{argmin}} \{ \|\mathbf{A}\mathbf{x} - \mathbf{b}\|_{W_\eta}^2 + \lambda^2 \|L(\mathbf{x} - \mathbf{x}_0)\|_2^2 \}$$

Mapping L defines basis for \mathbf{x} with prior \mathbf{x}_0

Weighting $W_\eta = C_\eta^{-1}$, $\|\mathbf{y}\|_{W_\eta} = \mathbf{y}^T W_\eta \mathbf{y}$. Whitens noise in \mathbf{b} .

Requires automatic estimation of λ^{opt}

Large Scale Problems use Iterative Solve (Notation)

LSQR Let $\beta_1 := \|\mathbf{b}\|_2$, and $\mathbf{e}_1^{(t+1)}$ first column of I_{t+1}
Generate, lower bidiagonal $B_t \in \mathcal{R}^{(t+1) \times t}$, column
orthonormal $H_{t+1} \in \mathcal{R}^{m \times (t+1)}$, $G_t \in \mathcal{R}^{n \times t}$

$$AG_t = H_{t+1}B_t, \quad \beta_1 H_{t+1}\mathbf{e}_1^{(t+1)} = \mathbf{b}.$$

Projected Problem on projected space:

$$\mathbf{w}_t(\zeta_t) = \operatorname{argmin}_{\mathbf{w} \in \mathcal{R}^t} \{ \|B_t \mathbf{w} - \beta_1 \mathbf{e}_1^{(t+1)}\|_2^2 + \zeta_t^2 \|\mathbf{w}\|_2^2 \}.$$

Projected Solution depends on ζ_t^{opt}

$$\mathbf{x}_t(\zeta_t^{\text{opt}}) = G_t \mathbf{w}_t(\zeta_t^{\text{opt}})$$

Generally: $\zeta_t^{\text{opt}} \neq \lambda^{\text{opt}}$

- (i) Determine optimal t The choice of the subspace impacts the regularizing properties of the iteration: For large t noise due to numerical precision and data error enters the projected space.
- (ii) Determine optimal ζ_t How do regularization parameter techniques translate to the projected problem?
- (iii) Relation optimal ζ_t and optimal λ Given t how well does optimal ζ_t for projected space yield optimal λ for full space, or when is this the case?

Residual: $\mathbf{R}^{\text{full}}(\mathbf{x}_t) = A\mathbf{x}_t - \mathbf{b}$.

Influence Matrix $A(\lambda) = A(A^T A + \lambda^2 I)^{-1} A^T$

UPRE : Full problem

$$\lambda^{\text{opt}} = \underset{\lambda}{\operatorname{argmin}} \{ \|\mathbf{R}^{\text{full}}(\mathbf{x}_t(\lambda))\|_2^2 + 2 \operatorname{Tr}(A(\lambda)) - m \} = \underset{\lambda}{\operatorname{argmin}} \{ U^{\text{full}}(\lambda) \}.$$

Using the projected solution for parameter λ and

$$\operatorname{Tr}((AG_t)(\lambda)) = \operatorname{Tr}(B_t(\lambda))$$

$$\begin{aligned} U^{\text{full}}(\lambda) &= \|((AG_t)(\lambda) - I_m)\mathbf{b}\|_2^2 + 2 \operatorname{Tr}((AG_t)(\lambda)) - m \\ &= \|\beta_1(B_t(\lambda) - I_{t+1})\mathbf{e}_1^{t+1}\|_2^2 + 2 \operatorname{Tr}(B_t(\lambda)) - m \end{aligned}$$

λ^{opt} for $U^{\text{full}}(\lambda)$ can be estimated for projected problem

Is λ^{opt} relevant to ζ_t^{opt} for the projected problem?

Noise in the right hand side For $\mathbf{b} = \mathbf{b}^{\text{true}} + \boldsymbol{\eta}$, $\boldsymbol{\eta} \sim \mathbb{N}(0, I_m)$

$$\beta_1 \mathbf{e}_1^{t+1} = H_{t+1}^T \mathbf{b} = H_{t+1}^T \mathbf{b}^{\text{true}} + H_{t+1}^T \boldsymbol{\eta}.$$

Noise in projected right hand side $\beta_1 \mathbf{e}_1^{t+1}$, satisfies

$$H_{t+1}^T \boldsymbol{\eta} \sim \mathbb{N}(0, I_{t+1})$$

Immediately

$$\begin{aligned} U^{\text{proj}}(\zeta_t) &= \|\beta_1 (B_t(\zeta_t) - I_{t+1}) \mathbf{e}_1^{(t+1)}\|_2^2 + 2 \text{Tr}(B_t(\zeta_t)) - (t+1) \\ &= U^{\text{full}}(\zeta_t) + m - (t+1). \end{aligned}$$

Minimizer of $U^{\text{proj}}(\zeta_t)$ is minimizer of $U^{\text{full}}(\zeta_t)$

ζ_t^{opt} calculated for projected problem may not yield λ^{opt} on full problem

ζ_t^{opt} depends on t , λ^{opt} depends on $m^* =: \min(m, n)$

Trace Relations By linearity and cycling.

$$\text{Tr}(A(\lambda)) = \text{Tr}(A(A^T A + \lambda^2 I_n)^{-1} A^T) = m^* - \lambda^2 \sum_{i=1}^{m^*} (\sigma_i^2 + \lambda^2)^{-1}$$

$$\text{Tr}(B_t(\zeta_t)) = t - \zeta_t^2 \sum_{i=1}^t (\gamma_i^2 + \zeta_t^2)^{-1}.$$

Approximate Singular Values IF $\sigma_i \approx \gamma_i$, $1 \leq i \leq t^* \leq t$,
 $\sigma_{t^*}^2 / (\sigma_{t^*}^2 + \lambda^2) \gg \sigma_i^2 / (\sigma_i^2 + \lambda^2) \approx 0$, $i > t^*$,

$$\text{Tr}(A(\lambda)) \approx \text{Tr}(B_{t^*}(\lambda)) + \sum_{i=t^*+1}^{m^*} \sigma_i^2 (\sigma_i^2 + \lambda^2)^{-1} \approx \text{Tr}(B_{t^*}(\lambda)).$$

If t^* approx numerical rank A , $\zeta_t^{\text{opt}} \approx \lambda^{\text{opt}}$ for $\mathcal{K}_{t^*}(A^T A, A^T \mathbf{b})$

GCV: [CNO08] *weighted* GCV is introduced for $\omega > 0$.

$$G^{\text{proj}}(\zeta_t, \omega) = \frac{\|\mathbf{R}^{\text{proj}}(\mathbf{w}_t(\zeta_t))\|_2^2}{(\text{Tr}(\omega B_t(\zeta_t) - I_{t+1}))^2}, \quad G(\lambda) = G^{\text{proj}}(\lambda, 1).$$

Optimal Analysing as for UPRE: $\omega = \frac{t+1}{m} < 1$.

Discrepancy Principle Seek λ such that

$$\|\mathbf{R}^{\text{full}}(\mathbf{x}(\lambda))\|_2^2 = \delta \approx m. \text{ To avoid over smoothing:}$$
$$\delta = vm, v > 1$$

Discrepancy for the Projected Problem Seek ζ_t such that

$$\|\mathbf{R}^{\text{proj}}(\mathbf{w}_t(\zeta_t))\|_2^2 \approx \delta^{\text{proj}} = v(t+1).$$

We do not obtain in these cases $\zeta_t^{\text{opt}} \approx \lambda^{\text{opt}}$

Noise revealing function: [HPS09] suppose θ_j and β_j on diagonal and sub diagonal of B_t

$$\rho(t) = \prod_{j=1}^t (\theta_j / \beta_{j+1})$$

Optimal t is given by (for user determined t^{\min})

$$t^{\text{opt}-\rho} = \min_{t > t^{\min}} \{ \operatorname{argmax}(\rho(t)) \} + \text{step}$$

step= 2 is to assure that noise has entered the entries in $\rho(t)$ and hence the basis.

t^{\min} is chosen based on examination of $\rho(t)$.

Only useful if discrete Picard condition holds [HPS09].

Identifying optimal subspace size t :

Minimization of the GCV for the truncated SVD of B_{t^*} [CKO15]

Projected subspace size is defined to be t^*

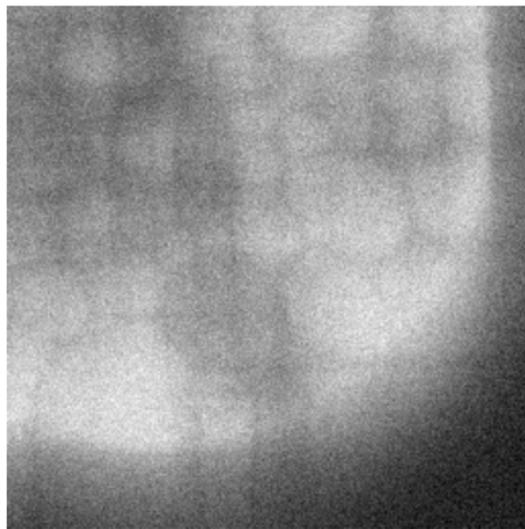
$$\mathcal{G}(t, t^*) = \frac{t^*}{(t^* - t)^2} \sum_{t+1}^{t^*} |\mathbf{u}_i^T \mathbf{b}|^2.$$

Optimal t is given by

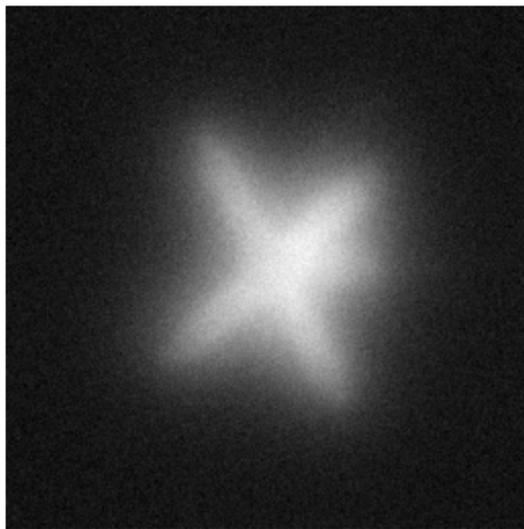
$$t^{\text{opt-G}} = \underset{t}{\operatorname{argmin}} \mathcal{G}(t, t^*)$$

Does not require Picard condition, but $t^{\text{opt-G}}$ depends on t^*

Application for Two Dimensional Examples



(a) Data



(b) Data

Figure: Data for grain and satellite images with blur and noise level 10%.

Noise Revealing Function $\rho(t)$: comparing $t^{\text{opt}-\rho}$, $t^{\text{opt}-\mathcal{G}}$, $t^{\text{opt}-\min}$

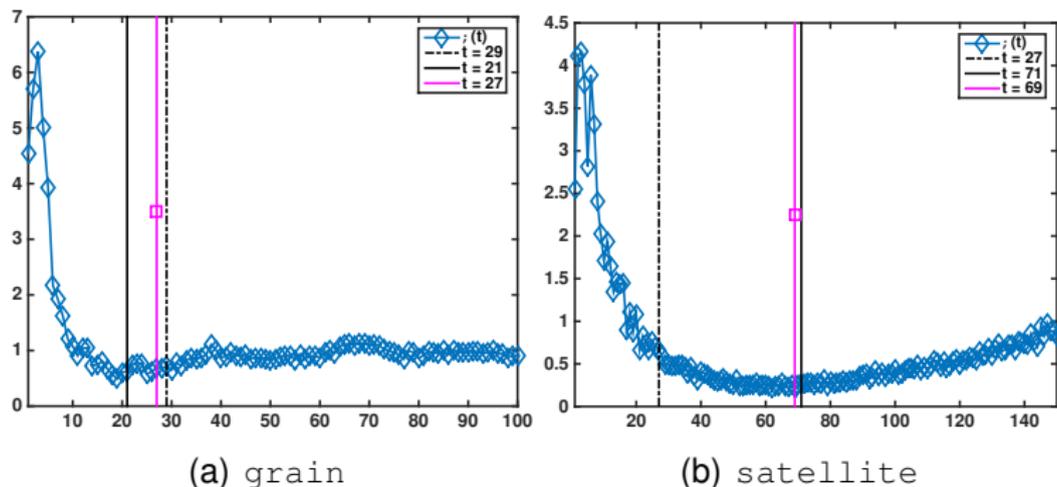
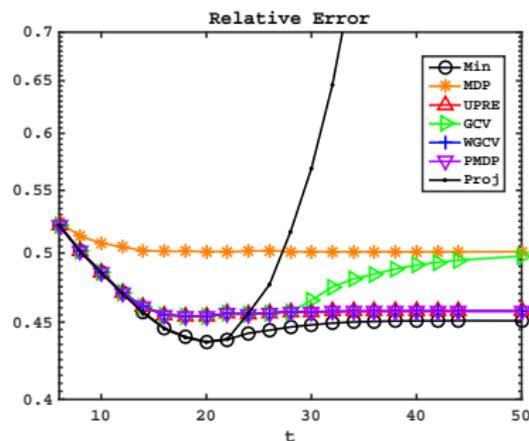
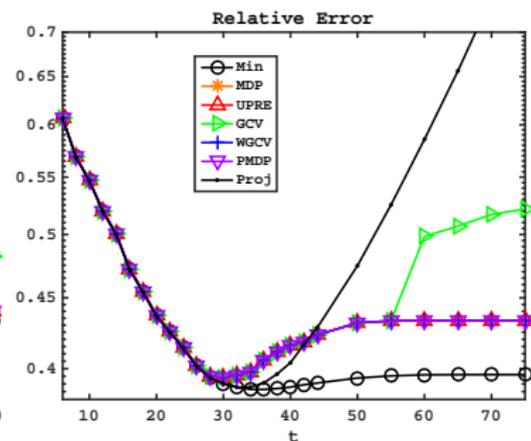


Figure: $\rho(t)$ using $t^{\min} = 25$. Dashed-dot $t^{\text{opt}-\rho}$, magenta $t^{\text{opt}-\mathcal{G}}$ and black $t^{\text{opt}-\min}$, location of minimum for $\rho(t)$ plus step.

Evaluating Image Quality : Relative error



(a) grain RE



(b) satellite RE

Figure: Relative error (RE) with increasing t . Solid line in each case is solution with projection and without regularization.

UPRE, WGCV and PMDP outperform GCV

Solutions for different t^{opt} : (MIN, $t^{\text{opt}-\min}$, $t^{\text{opt}-\mathcal{G}}$, $t^{\text{opt}-\rho}$) Noise level 10%

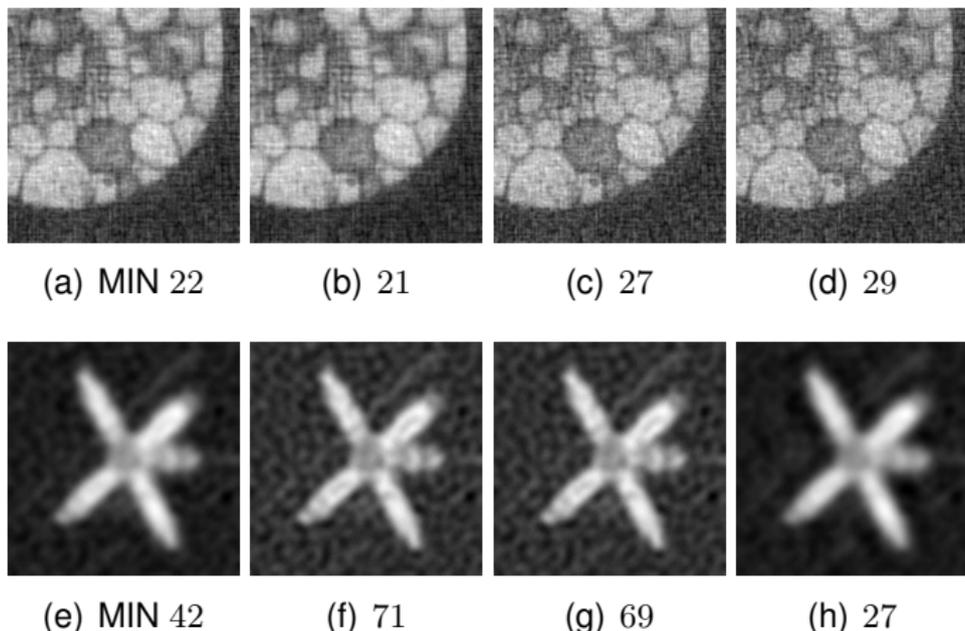


Figure: UPRE to find ζ . Solutions obtained for $t^{\text{opt}-\rho}$, $t^{\text{opt}-\min}$ and $t^{\text{opt}-\mathcal{G}}$ and MIN.

Solutions inadequate

Minimum Support Stabilizer Regularization operator $L^{(k)}$.

$$(L^{(k)})_{ii} = ((\mathbf{x}_i^{(k-1)} - \mathbf{x}_i^{(k-2)})^2 + \beta^2)^{-1/2} \quad \beta > 0$$

Parameter β ensures $L^{(k)}$ invertible

Invertibility use $(L^{(k)})^{-1}$ as right preconditioner for A

$$(L^{(k)})_{ii}^{-1} = ((\mathbf{x}_i^{(k-1)} - \mathbf{x}_i^{(k-2)})^2 + \beta^2)^{1/2} \quad \beta > 0$$

Initialization $L^{(0)} = I$, $\mathbf{x}^{(0)} = \mathbf{x}_0$. (might be 0)

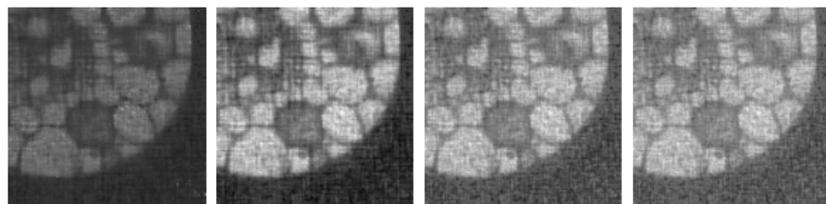
Reduced System When $\beta = 0$ and $\mathbf{x}_i^{(k-1)} = \mathbf{x}_i^{(k-2)}$ remove column i , \hat{A} is AL^{-1} with columns removed.

Update Equation Solve $\hat{A}\hat{\mathbf{y}} \approx \mathbf{R} = \mathbf{b} - A\mathbf{x}^{(k-1)}$. With correct indexing set $\mathbf{y}_i = \hat{\mathbf{y}}_i$ if updated, else $\mathbf{y}_i = 0$.

$$\mathbf{x}^{(k)} = \mathbf{x}^{(k-1)} + \mathbf{y}$$

Cost of $L^{(k)}$ is minimal

Solutions t^{opt} after two steps IRR: (MIN, $t^{\text{opt}-\min}$, $t^{\text{opt}-\mathcal{G}}$, $t^{\text{opt}-\rho}$)

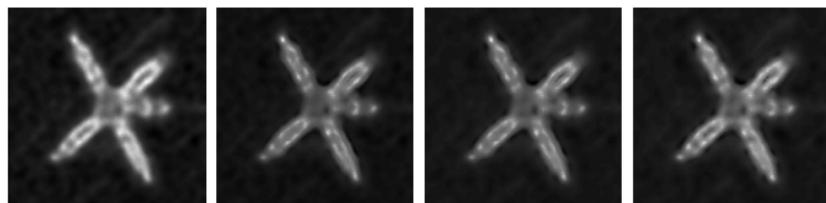


(a) 19 $k = 3$

(b) 21

(c) 27

(d) 29



(e) 35

(f) 71

(g) 69

(h) 27

Figure: IRR $k = 2$ Grain $k = 2$ MIN solution is at $t^{\text{opt}-\min}$, show $k = 3$.

Solutions are stabilized less dependent on t

Relative error with k : 5% error using UPRE

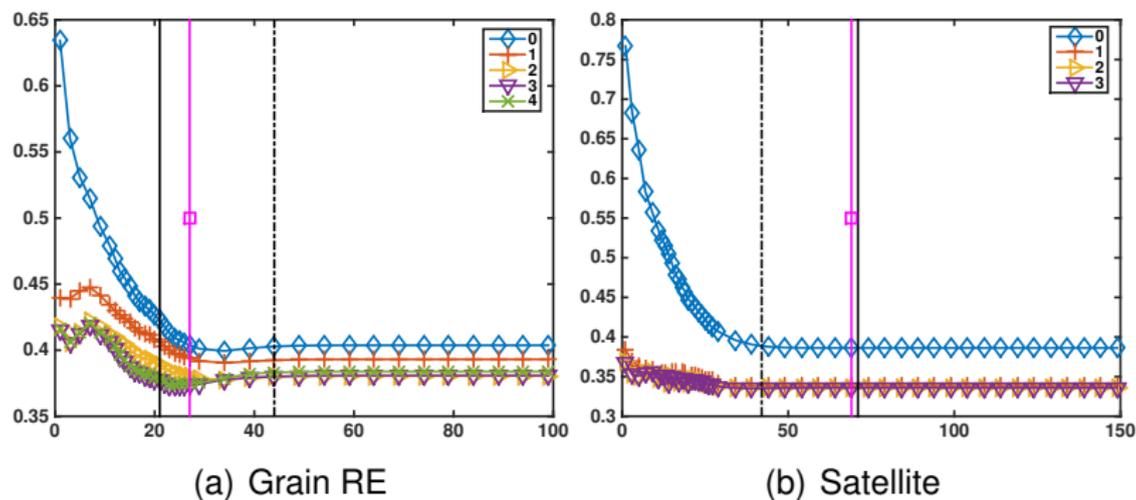


Figure: Relative errors decrease initially with k and then increase. Dashed-dot $t^{\text{opt}-\rho}$, magenta $t^{\text{opt}-\mathcal{G}}$, black $t^{\text{opt}-\min}$.

Noise revealing function $\rho(t)$ with k 5% error

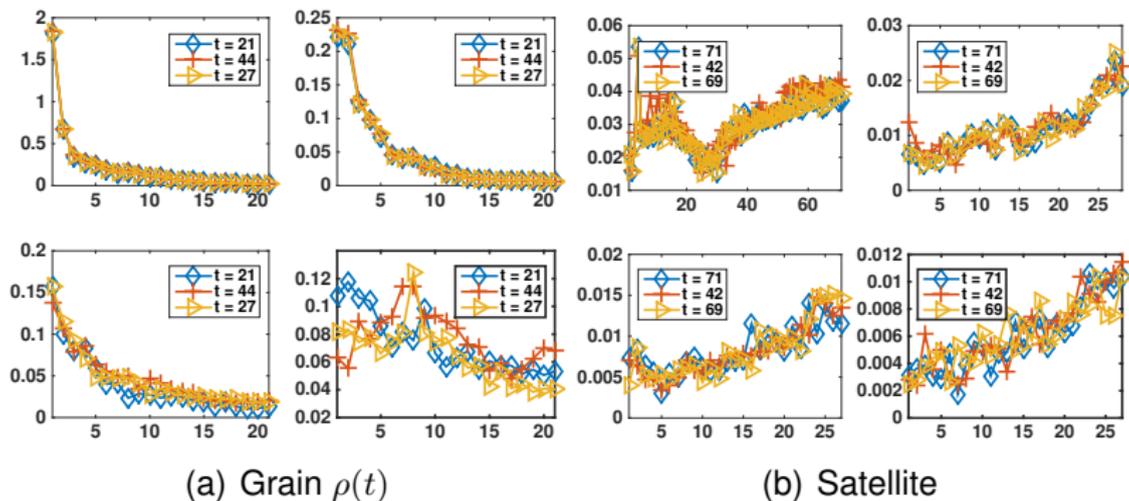


Figure: Determining t^{opt} with k for 5% noise using $\rho(t)$. Stopping Criteria : Grain $k = 4$ noise enters, use $k = 2$. Satellite $k = 3$ noise enters, use $k = 1$.

Solving the Gravity Inversion problem

Undersampled Gravity inversion $m = 4900, n = 98000$

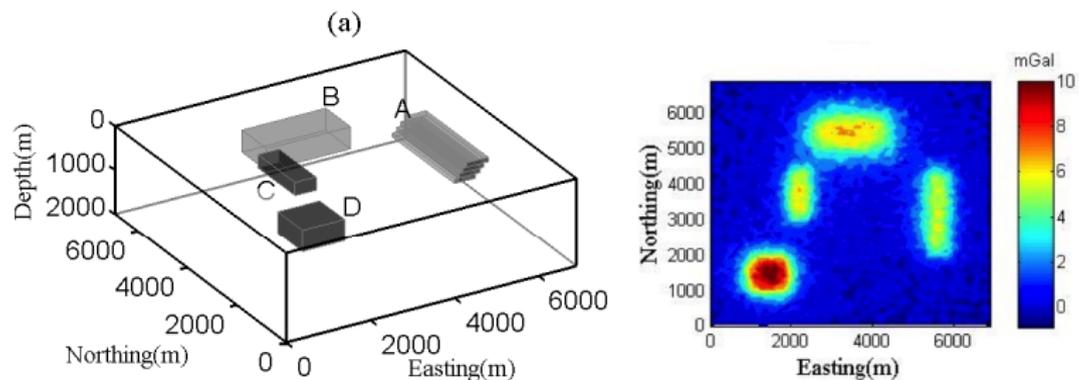


Figure: (a) The perspective view of the model. Four different bodies embedded in an homogeneous background. Densities of A and B are 0.8 g cm^{-3} and C and D are 1 g cm^{-3} ; (b) The noise contaminated gravity anomaly due to the model.

Undersampled Gravity inversion

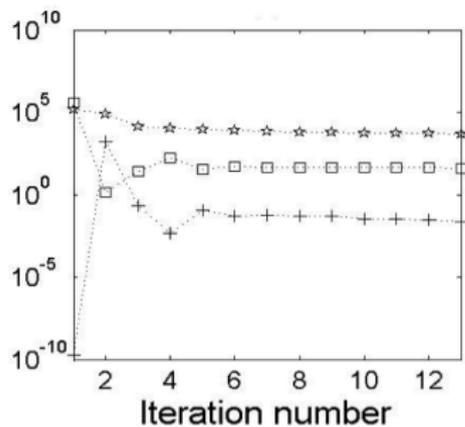
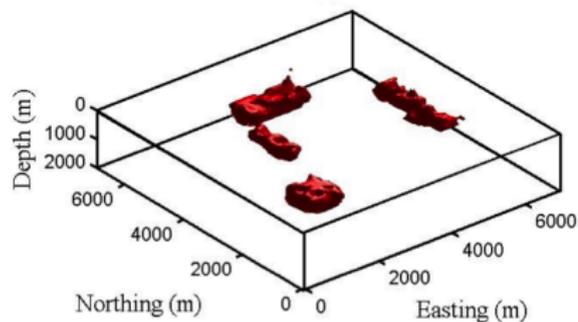


Figure: The reconstructed model with $t = 250$ and the L_1 stabilizer with $\beta^2 = 1.e-9$. Data misfit ★, the regularization term, +, regularization parameter □ with iteration

True Data:

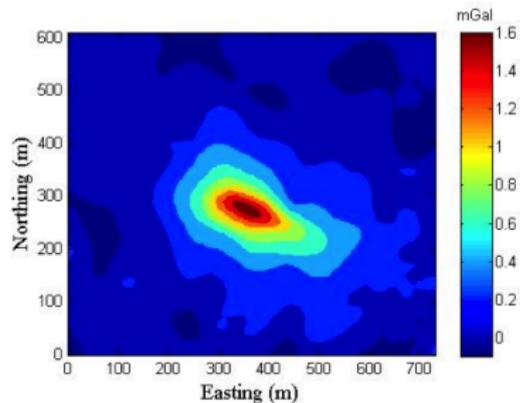


Figure: Residual Anomaly of Mobrun ore body, Noranda, Quebec, Canada.

Reconstructed Model

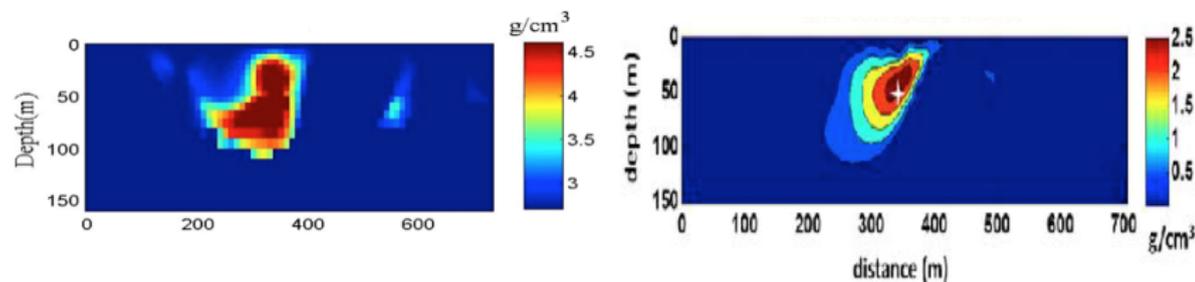


Figure: The reconstructed model with $t = 300$ and the L_1 stabilizer with $\beta^2 = 1.e-9$. (a) cross-section at $y = 285$ m and (b) comparison From Ialongo et.al (2014)

3-D perspective

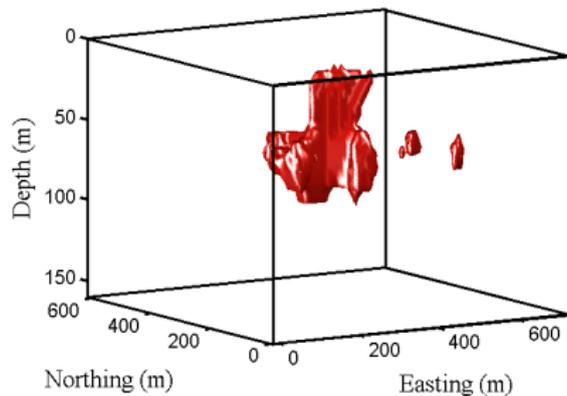


Figure: 3D view of the recovered model, the density cut off is 4 g cm^{-3} .

UPRE/WGCV regularization parameter estimation explained for projected problem.

ζ_t^{opt} , λ^{opt} related across levels

Underdetermined problems are also solved.

Iteratively Reweighted Regularization stabilizes the projected solution

Sensitivity to choice of t^{opt} reduced by IRR

t^{opt} can be estimated using $\rho(t)$, use $t^{\text{opt}-\min}$ as independent of other parameters

t^{opt} effectively determines a truncation of the SVD for B_t : use B_t and G_t , but truncated solution.

Some key references



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