



#### **Conceptual Climate Models**

Minitutorial Part I

Esther Widiasih University of Hawaii-West Oahu

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### Summary

- A brief history of Earth's climate
- What is climate?
- Some physics background
- OD Energy Balance Model (EBM)
- Tipping points/ Bifurcations/ Hysteresis
- 1D EBM (infinite dimensional)

#### A brief history of climate

Temperature of Planet Earth



Major features: glacial cycles with a change in amplitude and frequency, some warm climate in the past, hockey stick at PETM.

Image source: Wikipedia



• Climate = 30 year average of weather

Weather: Will I need an umbrella tomorrow?

Climate: Should I own an umbrella?

Slide credit: S. Oestreicher, JMM Minitutorial 2013

#### How do we observe climate?



http://cdiac.ornl.gov/d13\_flask\_mauna\_loa.html http://www.dartmouth.edu/~mpayres/People/Sharon.7506.web.jpg http:// www.whoi.edu/ooi\_cgsn/auvs-gliders?tid=1621&cid=137956&article=95673

#### How do we observe past climate?



#### Benthic foraminifera

Hold the long term memory of Earth's climate

Image source: usgs.gov

### How do we observe past climate?





#### Antartic ice core

Image source: http://earthobservatory.nasa.gov/Features/CarbonCycle/page4.php

#### How does CO<sup>2</sup> affect climate?



http://climate.nasa.gov/climate\_resources/24/

https://medium.com/@350/temperature-check-1f076624c55b#.mvvf6oqf8 https://www.ipcc.ch/publications\_and\_data/ar4/wg1/en/spmsspm-projections-of.html

### How to model climate?

 No detail is too small: Global Circulation Models GCM

 The rest is details: Conceptual Climate Models CCM

Slide source: S. Oestreicher JMM Minitutorial 2013.



Solar

Reflected Energy www.nasa.gov/vision/earth/lookingatearth/ice\_clouds.htm

Sun

### **Global Circulation Models**

## Complicated choices starting from the grids.





Must decide the the processes of each part and interactions among them.

Slide source: S. Oestreicher JMM Minitutorial 2013.

### **Global Climate Models**

- **Processes:** physics, biology, chemistry
- Computer Science Data mining, coupling nonsimilar grids, error analysis, parallel processing, time optimization
- Statistics

Extreme events, trends, and averaging

• Mathematics Data assimilation, numerical analysis, PDE



Slide source: S. Oestreicher (now S. Schumacher), JMM 2013

### A Simpler Life: Conceptual Climate Modeling

Can one develop a <u>mathematical</u> model that captures Earth's climate, ie through **temperature, ice cover, and CO2 level**?



Temperature of Planet Earth

Image source: Wikipedia

### The Modeling Cycle



### Some well known laws

- 1. Conservation of energy: Energy in = Energy out
- 2. Earth gets its energy from the sun
- 3. The Stefan Boltzmann black body radiation:

(Incoming solar energy is used to heat up the planet)

"A black body will emit a certain amount of energy *E* depending on it's temperature *T*."

#### $\boldsymbol{E} = \boldsymbol{\sigma}\boldsymbol{T}^4$

# What does Earth do with all that energy?



Image source: Trenberth, Kiehl, Fazulo http://scied.ucar.edu/radiation-budget-diagram-earth-atmosphere

The climate of Earth is represented by one point, the global annual average temperature.

#### **ENERGY BALANCE EQUATION**

#### Energy in = Energy out

Any left over energy is used to heat up the planet

**Temperature change = Energy in – Energy out** 



$$R\frac{dT}{dt} = Q - \sigma T^4$$
$$= Q - \sigma T^4$$

- Q = Solar energy received (343 Watt  $m^{-2}$ )
- R = heat capacity (unit W s m<sup>-2</sup> K<sup>-1</sup>)
- $\sigma$  = Stefan Boltzman constant ( 5.67x10<sup>-8</sup> W m <sup>-2</sup> K <sup>-4</sup>)
- T = Temperature (in Kelvin)

#### Exercise:

1. Verify the units

2. Assuming Earth's climate is in equilibrium,

what is Earth's global annual average temperature?



$$R\frac{dT}{dt} = Q - \sigma T^4$$
$$= Q - \sigma T^4$$

T<sub>equilibrium</sub> ~ **279K = 6°C** Observed Earth's global temperature is about **287K = 14°C** 





$$R\frac{dT}{dt} = Q - \sigma T^4$$
$$= Q - \sigma T^4$$

T<sub>equilibrium</sub> ~ **279K = 6°C** Observed Earth's global temperature is about **287K = 14°C** 

Add more process in the model?



### **Overview of climate**



Two types of Earth's reradiation:

Short wave and long wave.

Albedo/ reflectivity affects short wave reradiation.

### Some <u>not</u> so well known laws

4. Ice albedo feedback:ice/ snow –lighter color, reflects energyland/ ocean –darker color, absorbs energy



The Ice Albedo Feedback



Figure 1: The modeling cycle.

Q = 343 Watt m<sup>-2</sup>

- $\sigma$  = Stefan Boltzman constant ( 5.67x10<sup>-8</sup> W m <sup>-2</sup> K <sup>-4</sup>)
- α = planetary albedo ~ 0.3 (non dimensional constant)Exercise:

Modeling Earth's temperature

Zero spatial dimension

**Global Energy Balance Model** 

 $R\frac{dT}{dt} = Q - Q\alpha - \sigma T^4$ 

 $=Q(1-\alpha)-\sigma T^4$ 

1. Verify the units

2. Assuming that Earth's climate is in equilibrium, what is Earth's global annual average temperature?



$$R\frac{dT}{dt} = Q - Q\alpha - \sigma T^{4}$$
$$= Q(1 - \alpha) - \sigma T^{4}$$

Exercise 2.

Assuming Earth's climate is in equilibrium, what is Earth's global annual average temperature?

 $Q = 343 \text{ Watt } \text{m}^{-2}$ 

 $\sigma$  = Stefan Boltzman constant ( 5.67x10<sup>-8</sup> W m <sup>-2</sup> K <sup>-4</sup>)

 $\alpha$  = planetary albedo ~ 0.3 (non dimensional constant)



$$R\frac{dT}{dt} = Q - Q\alpha - \sigma T^4$$

$$=Q(1-\alpha)-\sigma T^4$$

2. Assuming Earth's climate is in equilibrium, what is Earth's global annual average temperature?

Q = 343 Watt m<sup>-2</sup>  $\sigma$  = Stefan Boltzman constant ( 5.67x10<sup>-8</sup> W m <sup>-2</sup> K <sup>-4</sup>)  $\alpha$  = planetary albedo ~ 0.3 (non dimensional constant) T<sub>eq</sub> = 255 K = -18°C!!! Far from the observed 14°C. The model gets worse.



### **Overview of climate**



Two types of Earth's reradiation:

Short wave and long wave.

Long wave radiation is due to eg. green house effect.



Modeling Earth's temperature Global Energy Balance Model

$$R\frac{dT}{dt} = Q(1-\alpha) - \varepsilon \cdot \sigma T^4$$

Stefan Boltzmann Law is radiation law for <u>black body</u>. Since Earth is not black body, must use different approximation of the long wave re-radiation.

Green house gas effect must be included: introduce the factor  $\epsilon$ .

Exercise 2b. Find the value of  $\varepsilon$  that fits the observation.



#### Modeling Earth's temperature Global Energy Balance Model

$$R\frac{dT}{dt} = Q(1-\alpha) - \varepsilon \cdot \sigma T^4$$

Stefan Boltzmann Law is radiation law for <u>black body</u>. Since Earth is not black body, must use different approximation of the long wave re-radiation.

Green house gas effect must be included: introduce the factor ε.

Exercise 2b. Find the value of  $\varepsilon$  that fits the observation. Answer:  $\varepsilon = 0.6$  will do it.

#### Modeling Earth's temperature **Global Energy Balance Model**

$$R\frac{dT}{dt} = Q(1-\alpha) - (A+BT)$$

Cycle # 3 Another

nnrnach

Stefan Boltzmann Law is radiation law for black body. Since Earth is not black body, must use different approximation of the long wave re-radiation.

Graves, et al (1993):  $\sigma T^4$  instead use A + B T  $A = 202 Wm^{-2} and B = 1.9 Wm^{-2}C^{-1}$ are constants obtained from satellite observation.

#### Modeling Earth's temperature profile One spatial dimension Zonal Energy Balance Model

Some assumptions that make the work easier:

- 1. Symmetry about the equator, so we only look at eg. the northern hemisphere.
- 2. Annual average temperature along the same latitude, say  $\theta$ .





### Modeling Earth's temperature profile Zonal Energy Balance Model



R = planetary heat capacity T = T(y) = T(y, t) s(y) = a distribution function  $\alpha(\eta, y) = \text{ the albedo at } y \text{ given that the ice line is at } \eta$   $\overline{T} = \int_{0}^{1} T(\xi) d\xi$  A = the "Greenhouse Gas Parameter" B, C are nonnegative parameters

$$\alpha(\eta, y) = \begin{cases} 0.32 & if \quad y < \eta \\ 0.47 & if \quad y = \eta \\ 0.62 & if \quad y > \eta \end{cases}$$



An example of the albedo function. Another is the smooth approximation of this.

#### Equilibrium temperature profile



#### Exercise:

Compute the equilibrium temperature profile for a fix  $\eta T_{\eta}^{*}(y)$ .

#### Equilibrium temperature profile





T<sup>\*</sup><sub>n</sub> (y).



### ICE LINE EVOLUTION

$$\frac{d\eta}{dt} = \varepsilon(T(\eta, t) - T_C)$$



The Ice Albedo Feedback

Compare the temperature at the ice line with a critical temperature  $T_c$ .

If too warm, ice melts, ice line moves to pole. If too cold, ice forms, ice line moves to equator.

#### ICE LINE EVOLUTION

$$\frac{d\eta}{dt} = \varepsilon(T(\eta, t) - T_C)$$





**The Ice Albedo Feedback** 

#### Low latitude initial ice line

#### More simulations



#### Mid latitude initial ice line

High latitude initial ice line

#### Ice line stability?



#### NEXT?

- Reduction of the infinite dimensional system to finite dimensional system.
- Bifurcation analysis

#### **Resources:**

Modeling module http://dimacs.rutgers.edu/MPE/Energy/DIMACS-EBM.pdf

Conceptual Climate Module

https://mcrn.hubzero.org/resources/523/supportingdocs. Material developed for the MAA-NCS Summer Seminar Conceptual Climate Models, held in Minneapolis, July 2013. Contributors: A. Barry, R. McGehee, S. Oestreicher, J.A. Walsh, E. Widiasih.

## What little bugs in the ocean **don't** tell us: Snowball Earth

Snowball Earth http://www.wwnorton.com/college/geo/animations/ snowball\_earth.htm

#### What the model says about Snowball Earth



Image source: Wikipedia

#### Did the snowball earth ever happen?



Some evidence: Dropstone deposited by glacier into a marine sediment (Namibia)

> Image source: Terra Nova, 2002